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# Coarse- versus fine-grain quartz OSL and cosmogenic <sup>10</sup>Be dating of deformed fluvial terraces on the northeast Pamir margin, northwest China



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## ABSTRACT

Along the NE Pamir margin, flights of late Quaternary fluvial terraces span actively deforming fault-related folds. We present detailed results on two terraces dated using optically stimulated luminescence (OSL) and cosmogenic radionuclide <sup>10</sup>Be (CRN) techniques. Quartz OSL dating of two different grain sizes (4-11 µm and 90-180 µm) revealed the fine-grain quartz fraction may overestimate the terrace ages by up to a factor of ten. Two-mm, small-aliquot, coarse-grain quartz OSL ages, calculated using the minimum age model, yielded stratigraphically consistent ages within error and dated times of terrace deposition to  $\sim 9$  and  $\sim 16$  ka. We speculate that, in this arid environment, fine-grain samples can be transported and deposited in single, turbid, and (sometimes) nighttime floods that prevent thorough bleaching and, thereby, can lead to relatively large residual OSL signals. In contrast, sand in the fluvial system is likely to have a much longer residence time during transport, thereby providing greater opportunities for thorough bleaching. CRN <sup>10</sup>Be depth profiles date the timing of terrace abandonment to ~8 and ~14 ka: ages that generally agree with the coarse-grain quartz OSL ages. Our new terrace age of  $\sim$  13–14 ka is broadly consistent with other terraces in the region that indicate terrace deposition and subsequent abandonment occurred primarily during glacial-interglacial transitions, thereby suggesting a climatic control on the formation of these terraces on the margins of the Tarim Basin. Furthermore, tectonic shortening rates calculated from these deformed terraces range from  $\sim 1.2$  to  $\sim 4.6$  mm/a and, when combined with shortening rates from other structures in the region, illuminate the late Quaternary basinward migration of deformation to faults and folds along the Pamir-Tian Shan collisional interface.

## 1. Introduction

Fluvial terraces can be excellent geomorphic markers that record recent tectonic deformation and uplift, in addition to capturing a landscape's response to climatic events, e.g., Repka et al. (1997); Pan et al. (2003); and Burbank and Anderson (2011). Fluvial terraces may form in response to changes in deformation, e.g., uplift rates, or climate, e.g., sediment supply or discharge, allowing a river to incise into fill or bedrock, thereby abandoning the previous riverbed. Whereas the original terrace surface mimics the gradient of the previous riverbed, it can be subsequently deformed by faults and folds. The deviation from the original terrace gradient can then be used to calculate deformation rates, if the terrace can be accurately dated. Thus, reliable dating of fluvial terraces is a key step to unraveling the geomorphic record and characterizing deformation rates. Dating has proven challenging in semi-arid to arid settings, where readily dateable materials, such as organic debris for <sup>14</sup>C dating, are uncommon. Recent advances in cosmogenic radionuclide (CRN), e.g., Gosse and Phillips (2001), and optically stimulated luminescence (OSL) dating (Aitken, 1998) have allowed successful dating of terrace surfaces in these settings, e.g., Rittenour (2008); Porat et al. (2009); Guralnik et al. (2011); Viveen et al. (2012); and Li et al. (2013). Given uncertainties associated with individual dating methods, dating of terraces using both cosmogenic <sup>10</sup>Be and OSL techniques has become increasingly common (Hetzel et al., 2004; Owen et al., 2006; DeLong and Arnold, 2007; Nissen et al., 2009; Fruchter et al., 2011; Owen et al., 2011; Guralnik et al., 2011; Lee et al., 2011; Viveen et al., 2012). Importantly, OSL and <sup>10</sup>Be date different geomorphic events: OSL dates the deposition of the sediments that aggrade above a strath, whereas cosmogenic <sup>10</sup>Be dates the abandonment and stabilization of the terrace surface. Thus, when combined

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Fig. 1. (A) Topography of the Pamir and surrounding area. (B) Simplified geologic map of the western Tarim Basin. Faults: MPT – Main Pamir Thrust, MST – Mingyaole South Thrust, MT – Mayikake Thrust, PFT – Pamir Frontal Thrust, TT – Takegai Thrust. Rivers: Bie – Bieertuokuoyi River, K – Kangsu River, Ka – Kalangoulvke River, Ke – Kezilesu River, Mar – Markansu River. (C) GoogleEarth image showing regional river system, <sup>10</sup>Be depth profile sample locations, and their respective upstream catchments.

with additional field data and observations, the difference between cosmogenic <sup>10</sup>Be- and OSL-derived ages may reveal important information about the geomorphic system and the formation of the terrace surfaces (Guralnik et al., 2011), such as the terrace aggradation rates, uplift and incision rates, and paleo-erosion rates.

Flights of fluvial terraces span large areas of the Pamir and Tian Shan foreland basins in the western Tarim Basin in northwest China (Fig. 1) (Bufe et al., 2016). Many of these terraces are deformed by active faults and folds and record the late Quaternary deformation of these structures (Scharer et al., 2006; Heermance et al., 2008; Li et al., 2012, 2017, 2013, 2015a, 2015b; Thompson Jobe et al., 2017). Deformation by numerous active structures in this area inhibits reliable correlation of undated fluvial terraces for calculating slip rates on faults and folds. Hence, to determine local deformation rates, terraces crossing each individual structure must be dated. Previously, several deformed terraces have been dated using fine-grain quartz OSL (Li et al., 2012, 2013; Thompson Jobe et al., 2017). Given that numerous younger terraces yielded ages similar to older, higher surfaces dated using the same grain size and protocol, we suspect that many of these samples from younger terraces may have been poorly bleached, resulting in ages that overestimate the actual depositional ages. To test this possibility, we compared both fine- and coarse-grain quartz OSL dating techniques for consistency with each other and with independent ages from cosmogenic <sup>10</sup>Be depth profiles.

Here, we present a case study of two fluvial terraces in the western Tarim Basin, focusing on (i) the applicability of quartz OSL and cosmogenic <sup>10</sup>Be dating techniques to date fluvial terraces in an arid,

tectonically active region of NW China; (ii) assessment of which grain sizes are most appropriate for both OSL and cosmogenic <sup>10</sup>Be dating; (iii) evaluation of which OSL age model is most appropriate for determining a reliable depositional age for these terraces and (iv) calculation of deformation rates on three active structures that deform these two terraces. Ultimately, our goal is to develop a regional chronology of the widespread terraces to assess slip rates and climate-tectonic interactions across the NE margin of the Pamir orogen. This study presents a subset of our data toward that end.

## 2. Regional setting

In NW China, the western Tarim Basin lies between the Pamir and Tian Shan at the northwestern end of the Himalayan-Tibetan orogen, which formed as a result of the Indo-Eurasian collision (Fig. 1). In the western Tarim Basin, numerous Miocene-to-Recent faults and folds deform the Cenozoic sedimentary basin fill along the margins of the basin (Chen et al., 2002; Scharer et al., 2006; Heermance et al., 2008; Thompson et al., 2015). The basin-bounding faults have also uplifted and exposed Mesozoic and Paleozoic sedimentary and metamorphic units that have served as source areas for Quaternary sediment within the Pamir and Tian Shan orogens (Sobel et al., 2013).

The Main Pamir Thrust, defining the northern margin of the Pamir Plateau, initiated approximately 20 Ma (Sobel and Dumitru, 1997), and deformation propagated basinward during the Late Miocene to form the Takegai and Pamir Frontal Thrusts (Thompson et al., 2015). During the Quaternary, deformation has been accommodated on the Pamir

Frontal, Takegai, and Main Pamir Thrusts along the margins of the Tarim Basin, although in the last 125 ka, most deformation has been focused along a narrow corridor between the Pamir and Tian Shan (Li et al., 2012, 2015b; Thompson Jobe et al., 2017). Since  $\sim 0.35$  Ma, the Pamir Frontal Thrust has maintained a nearly uniform shortening rate of 6–8 mm/a (Li et al., 2012).

The Kezilesu River (Fig. 1B and C) is the largest river in the region, currently trapped between the Pamir and Tian Shan and flowing eastward parallel to the regional structural trend. In the western Tarim Basin, all major tributaries flow either north from the Pamir or south from the Tian Shan and join the Kezilesu River (Fig. 1B and C). In the Tian Shan, the bedrock in the source area comprises older formations of carbonates, clastics, and some igneous rocks. In the Pamir, igneous and clastic units dominate the source area. The western Tarim Basin has a present-day arid to semi-arid climate with highly seasonal variations in precipitation, which influence fluvial discharge dynamics. Driven by snowmelt, intense storms, and cloudbursts, the highest flows occur in the spring and summer, whereas slower flows occur in the winter. These variations in seasonal discharge have implications for the transport of sediment through the fluvial system, i.e., flash floods, transient storage in overbank deposits, or on bars in between high discharge flows, and they affect the bleaching of grains used in OSL dating and the deposition of sediment on the landscape. Within the foreland, most rivers have beveled the underlying Tertiary strata and created suites of gravelcovered strath terraces during the late Quaternary (Scharer et al., 2006; Heermance et al., 2008; Li et al., 2012, 2013, 2015a, 2015b; Bufe et al., 2016, 2017; Thompson Jobe et al., 2017).

We selected two representative sites on fluvial terraces along the Pamir Frontal Thrust on the NE Pamir margin in the western Tarim Basin (Fig. 1 and Fig. 2), and sampled the sand-and-gravel cover above

the bedrock strath. At each site, we collected four OSL samples and one cosmogenic <sup>10</sup>Be depth profile of both sand and pebbles, resulting in a total of 8 OSL samples and 2 cosmogenic <sup>10</sup>Be depth profiles that comprise 12 sand samples and 4 pebble samples.

# 3. Study area

### 3.1. Mayikake

The Mayikake site sits within a large, flat aggradation surface, 15 by 10 km, near the village of Mavikake (Fig. 1B and Fig. 2A). The extensive terrace surface is bounded on its northern side by the Kezilesu River and on the other margins by Cenozoic bedrock, exposed by uplift on the Pamir Frontal Thrust (PFT). On the SW margin, the terrace surface is cut by the Bieertuokuovi Frontal Thrust, a segment of the PFT (Fig. 2A). This fault offsets both Holocene alluvial fans at the mountain front and river terraces formed by the Bieertuokuoyi River. SW-dipping Paleogene sediments were thrust over the fluvial terrace deposits along a fault dipping 75  $\pm$  5° SW. Striae on the surface of the fault plane have a rake of 34°, indicating a strike-slip to dip-slip ratio of ~3:2 (Li et al., 2012). Differential GPS surveys of the terrace surface offset by the Bieertuokuoyi Frontal Thrust indicate a vertical separation of ~38 m (Li et al., 2012). On the northern part of the terrace, the Mayikake Thrust, a gently, north-dipping fault, cuts the surface to produce a fault scarp that extends no longer than  $\sim 8 \text{ km}$  and is  $\sim 15 \text{ m}$  high (Li et al., 2012). Recent incision of the terrace surface crossing the fault plane reveals a dip of  $\sim 16^{\circ}$  (Li et al., 2012).

Fluvial terrace gravels, ranging from 2- to > 10-m thick, cap a strath eroded into lithified Cenozoic units. Since abandonment of the terrace surface, the river has incised between 20 and 70 m into the underlying



Fig. 2. (A) Simplified geologic map of the Mayikake basin (location in Fig. 1B). (B) Sample LED 11–355 (C) <sup>10</sup>Be depth profile pit on Mayikake terrace surface (D) Simplified geologic map of the southern Mingyaole anticline (D) Sample LED 11–357 (F) Sample LED 11–360. BFT – Bieertuokuoyi Frontal Thrust, MT – Mayikake Thrust, PFT – Pamir Frontal Thrust, N – Neogene, E – Paleogene, N2-Q – Miocene-Pleistocene conglomerates. Rivers: Bie – Bieertuokuoyi River, K – Kangsu River, Ka – Kalangoulvke River, Ke – Kezilesu River.

Location and lithology of OSL samples.

Sample No.	Latitude/Longitude	Elevation (m)	Terrace Level	Lithology	Depth (m)	Thickness of silt lens (cm) <sup>a</sup>
Mayikake						
LED11-210	39.5811°N/75.0957°E	1869	T2	muddy silt	1.9	8
LED11-209	39.5811°N/75.0957°E	1869	T2	muddy silt	2.2	15
LED11-355	39.5833°N/75.1043°E	1869	T2	silty fine sand	2.7	8
LED11-356	39.5833°N/75.1043°E	1869	T2	muddy silt	3.1	8
Mingyaole						
LED11-360	39.4835°N/75.3822°E	1670	T2	fine sand	0.6	8
LED11-357	39.4836°N/75.3821°E	1670	T2	silty fine sand	0.7	30
LED11-359	39.4835°N/75.3822°E	1670	T2	sandy silt	0.8	45
LED11-358	39.4836°N/75.3821°E	1670	T2	sandy silt	1.0	25

<sup>a</sup> The thickness of the lens from which sample was collected, centered on the depth of the sample.

bedrock. The terrace surface displays a poorly-to-moderately developed desert pavement, with fractured and highly weathered clasts. Imbricated pebble-cobble clasts, with uncommon interbedded laminated and cross-bedded sands and silts, compose the gravel cover. We collected four OSL samples from trenches on the hanging wall of the Mavikake Thrust. All samples were collected from 8- to 15-cm-thick, muddy silt and sandy silt lenses from depths of 1.9-3.1 m (Fig. 2B, Table 1). Although the sampling site is near a fault, we carefully selected sites within the hanging wall of the fault where the primary stratigraphy was preserved, such that we are confident that fault-related deposits, such as colluvial wedges or ponding against the fault scarp, were excluded. The cosmogenic<sup>10</sup>Be depth profile was collected from a separate, hand-dug pit located  $\sim 40$  m to the north of the fault scarp, where the surface had not been recently modified (Fig. 2C). Observations from the depth profile pit indicate no evidence of buried soils, depositional hiatuses, or extensive bioturbation.

Previous fine-grain quartz OSL dating of this terrace yielded an age of ~18.4  $\pm$  4.3 ka (20): an age correlated to a regional terrace level dated to the Last Glacial Maximum (LGM) (Li et al., 2012). This age, combined with the measured offsets, defined slip rates of ~3.6 and ~3.1 mm/a for the Pamir Frontal Thrust (Bieertuokuoyi Frontal Thrust in Li et al., 2012) and the Mayikake Thrust, respectively.

#### 3.2. Mingyaole

The second site is a fluvial terrace on the southern side of the growing Mingyaole fold, adjacent to the Kezilesu River (Fig. 2D). The Mingyaole anticline initiated ~1.6 Ma (Chen et al., 2005; Thompson, 2013), and has accommodated ~1.5 km of shortening at a mean rate of ~0.9 mm/a (Chen et al., 2005). As the anticline continued to grow, terrace surfaces on the flanks of the anticline have been deformed to produce a series of fold scarps and flexural-slip fault scarps (Chen et al., 2005, 2007; Li et al., 2015a, b). We sampled from the upper tread of the fold scarp on the T2 terrace. The fold scarp on the T2 surface is ~16 m high and dips ~25° to the south. The cumulative shortening absorbed by the fold scarp is ~10.1<sup>+1.9</sup>/<sub>-1.6</sub> m since the abandonment of the terrace surface (Li et al., 2015b). Despite the proximity to these tectonic features, intact sedimentary structures and field observations suggest a stratigraphy that is undisturbed by tectonic deformation.

Fluvial gravels ~5 m thick rest above a strath terrace that was beveled into Neogene sedimentary formations. Currently, the river flows ~60 m below the terrace surface and has incised through the terrace fill and underlying Neogene bedrock. The terrace surface has a poorly-to-moderately developed desert pavement, with fractured, weathered, and varnished clasts. The terrace deposits consist mostly of imbricated pebble-cobble clasts, with interbedded laminated and cross-bedded silty-sand lenses and uncommon massive, muddy silt beds ~ 10-cm thick (Fig. 2E and F). (See Li et al. (2015a) for a detailed description and geomorphic map of the terraces near Mingyaole.) We exploited man-made pits that were hand-dug ~ 2.5 m into the terrace surface.

Because the pits were not present when we previously visited this site in 2010 (samples were collected in 2011), we know the pits were dug less than a year before we sampled from them. We collected two OSL samples from each of two pits and the <sup>10</sup>Be depth profile from a third pit: all less than ~50 m apart. All OSL samples were collected from 7.5-to 45-cm-thick silty sand lenses from depths of 0.6–1 m (Fig. 2, Table 1). Observations from the depth-profile pit show no evidence of buried soils, depositional hiatuses, or extensive bioturbation. This terrace surface has not been previously dated, and our previous work suggested it might be equivalent to the LGM terrace at Mayikake based on surface characteristics.

## 4. OSL dating

OSL dating relies on the assumption that grains must be exposed to light for a certain length of time to remove the previous luminescence signal (Aitken, 1998). With insufficient exposure, grains remain incompletely bleached and retain a residual equivalent dose. Several factors affect the probability of bleaching sediments during fluvial transport, such as grain size (e.g., Stokes et al., 2001), turbidity (e.g., Berger and Luternauer, 1987), sediment load, water depth (e.g., Berger, 1990), sediment-transport distance (e.g., Stokes et al., 2001), and time in the fluvial system (Rittenour, 2008; Rhodes, 2011; ). Recent modeling and dating research suggest the opportunities for bleaching of the grains are largely related to the mode of transport through the fluvial system (Rittenour, 2008; Gray and Mahan, 2014; Cunningham et al., 2015). Because hydraulic conditions dictate that grains of different sizes will commonly have contrasting transport histories and may end up in different fluvial deposits, grains of different sizes are likely to have experienced different bleaching histories (Stokes et al., 2001; Rittenour, 2008). Beyond the tectonic implications, one goal of our research was to test the consistency of OSL dates from fine- and coarsegrain deposits in this arid setting.

## 4.1. OSL sample collection and analysis

We collected samples by hammering a metal tube parallel into the sediment layers. After removal from the surrounding sediment, we sealed the samples at both ends to prevent water loss and exposure to light during transport. The metal tubes were opened and the samples were processed under subdued red light at the Research Laboratory of Luminescence Dating at the Institute of Geology, China Earthquake Administration, in Beijing. All grain-size fractions were pretreated with 30%  $H_2O_2$  and 30% HCl to remove organics and carbonates, respectively. The fine-grain fraction (4–11 µm) was separated using Stokes' Law. The polymineralic fine-silt grains were immersed in hydrofluosilicic acid (40%) for three days in a centrifuge tube to isolate the quartz. The fine quartz grains were mounted on 9.7-mm steel discs from suspension in acetone. Coarse-grain samples (90–180 µm) were immersed in a 10% HF solution for 10 min, followed by a 40% HF bath for

40 min, then by 30% HCl for 40 min. The coarse-grain quartz grains were mounted on 9.7-mm steel discs using silicone gel to create small aliquots ( $\sim$ 2-mm mask diameter). The purity of the quartz was checked by IR stimulation and verified through observation of background IR signal and the typical 110° TL peak. Nevertheless, we performed an OSL-IR depletion test (Duller, 2003) on every coarse-grain aliquot.

## 4.2. OSL equipment and measurements

All fine-grain quartz samples, as well as coarse-grain quartz samples LED 11-210, LED 11-355, and LED 11-356, were measured using a Daybreak 2200 automated OSL reader, equipped with a combined blue  $(470 \pm 5 \text{ nm})$  and infrared  $(880 \pm 80 \text{ nm})$  LED OSL unit with a calibrated  ${}^{90}$ Sr/ ${}^{90}$ Y beta-radiation source (dose rate: 0.0327 Gy/s). Detection of the signal was through a 7-mm-thick U-340 glass filter. The coarse-grain samples LED 11-360 and LED 11-357, and additional LED 11-356 aliquots, in addition to preheat plateau and dose-recovery tests (Wintle and Murray, 2006) on sample LED 11-210, were measured using a Riso Reader model TL/OSL-DA-20 equipped with a calibrated  $^{90}$ Sr/ $^{90}$ Y beta-radiation sources (dose rate: 0.1051 Gy/s), blue (470  $\pm$  30 nm;  $\sim 50~mW/cm^2)$  and infrared (IR: 880 nm  $\pm$  80 nm,  $\sim$  145 mW/cm<sup>2</sup>) LEDs, and detection through a 7-mm-thick U-340 glass filter (Botter-Jensen et al., 2000). We conducted quality tests to ensure both OSL readers were returning comparable results and checked against independent <sup>14</sup>C dating results (Liu et al., 2010). We are convinced that using two different readers has not introduced any additional uncertainties beyond the uncertainties in the equivalent dose, hereafter noted by De. All luminescence measurements were made at 125 °C to prevent re-trapping in the 110 °C TL trap with both IR- and blue-light stimulation power at 80%. All  $D_e$  measurements were made using the sensitivity-corrected, multiple aliquot regenerative (SMAR) protocol for silt-sized quartz (Lu et al., 2007; Table S1), or a modified, single-aliquot regenerative (SAR) protocol for fine-sand quartz (Murray and Wintle, 2000; Table S2), with a thermal wash at 280 °C at the end of each cycle. On one sample, LED 11-210, we performed (1) dose-recovery tests to evaluate the ability to recover a known laboratory dose and (2) preheat plateau tests to evaluate any dependence on temperature (Wintle and Murray, 2006). Both of the tests were conducted using three aliquots for each temperature step between 180° and 260 °C. The preheat test identified a plateau between 220° and 280 °C, and the protocol was able to recover a regenerated dose within 10% of unity at 260 °C (see supplementary material, Fig. S1). Based on these results, we chose to apply a preheat temperature of 260 °C and a cut-heat temperature of 220 °C.

### 4.3. OSL data analysis

We used early background subtraction to calculate the  $D_e$  for all samples, i.e., the sum of the photons detected in the first 0.4 s or the first 0.23 s of the OSL decay curve for the Daybreak 2200 and Riso Readers, respectively, minus the sum of the next 1 s or 0.58 s, respectively. Given that some of the samples exhibited a medium component, use of early background subtraction isolates the fast component of the quartz.

Only aliquots (sub-samples) that satisfied the following criteria were used in the  $D_e$  calculation for small aliquot: (1) the OSL-IR depletion ratio was between 0.9 and 1.1, such that relative to OSL, no infrared signal exists above the background level (a response to infrared stimulation might signify contamination of the signal from feldspar) (Fig. 3C and F); (2) the recycling ratio was between 0.8 and 1.2 (Fig. S1C); (3) the recuperated OSL signal was less than 5% of the natural signal. Regenerative doses were fit with a saturating exponential equation to calculate the growth curve, or dose-response curve. The natural signal from each sample (fine-grain quartz) or aliquot (coarsegrain quartz) was then used to calculate the  $D_e$ .

#### 4.4. OSL dose-rate calculation

To calculate the dose rate for most samples, we used  $\sim 100 \text{ g}$  of sediment from the surrounding 30 cm of sediment. However, for one sample (Mayikake), the dose rate was only calculated using sediment in the tube. We quantified the concentrations of U, Th, K, and Rb using an ICP-MS, as well as the water content and cosmic ray contribution in each sample, with the exception of LED 11-356. For these samples, we measured the total alpha counts (Table 2) to calculate a bulk alpha rate from U and Th following the conversion factors in Aitken (1985) and determined the concentration of K. Elemental concentrations were measured at the ALS mineral lab in Reno. NV. Alpha counts were measured using 583 Davbreak alpha counters at the Institute of Geology, China Earthquake Administration, in Beijing, China. We note that both methods yielded relatively consistent results, with little variation in the dose rates across the region (see Table 2). An alpha efficiency of 0.04  $\pm$  0.02 for silt (4–11 µm) quartz (Rees-Jones, 1995) was used for the fine-grain dose-rate calculation. The cosmic-ray dose rate was calculated following Prescott and Hutton (1994).

The natural and saturated water content was measured in the laboratory. The possibility of time-varying water content was considered in more detail to calculate an average total dose-rate. We assumed the sediments were water-saturated when they were initially deposited. Because of the modern arid climate, the samples were nearly dry when we collected them, but we do not know at what point since initial deposition the samples were raised above the water table. Therefore, we consider a water content that varies from a completely dry sample (0% water content) to one with a saturated water content, by dividing the saturated water content in half and assigning a 100% error (Li et al., 2012, 2013). This range should cover all likely values.

## 4.5. Choice of OSL age model

Several statistical procedures exist to determine the appropriate model to calculate the paleodose of incompletely bleached samples, e.g., Bailey and Arnold (2006), for single-grain and small-aliquot data. The central age model for normally distributed data (CAM: Galbraith et al., 1999) and minimum age model for skewed and scattered data (MAM: Galbraith et al., 1999) are the models most commonly applied to fluvial deposits. Other studies, e.g., Roberts et al. (2000) and Rodnight et al. (2006), have also employed the finite mixture model (FMM: Galbraith and Green, 1990) to determine a burial dose for heterogeneously bleached samples with discrete dose populations. For a recent review of statistical methods, see Galbraith and Roberts (2012) and Kunz et al. (2014).

The  $D_e$  distributions of all coarse-grain quartz samples are skewed, with a tail of higher  $D_e$  values, yet no negative  $D_e$  values. We follow the statistical procedures outlined in previous studies (Olley et al., 2004; Bailey and Arnold, 2006) to guide our choice of age model and determine an age for each terrace surface. Thus, to choose the appropriate age model, we relied on the degree of over-dispersion, absence of negative  $D_e$  values, skewness (Bailey and Arnold, 2006), and geomorphic context. OSL ages were calculated using the R Luminescence software package (Kreutzer et al., 2012) and Excel spreadsheets. OSL ages are presented with 1-standard-error.

## 5. Cosmogenic <sup>10</sup>Be depth profile sampling & analysis

Cosmogenic-nuclide depth profiling relies on the predictable decrease of the <sup>10</sup>Be concentration with depth below the surface. As long as initial aggradation was rapid compared to the age of the terrace and the surface has been stable since abandonment, a depth-dependent trend in the <sup>10</sup>Be concentration can be used to date the terrace surface and determine the inheritance (Anderson et al., 1996; Repka et al., 1997). We collected cosmogenic depth profiles from ~2-m-deep pits located on unmodified fluvial terrace surfaces (Anderson et al., 1996;



Fig. 3. (A) Dose-response curve for LED 11–355, following SMAR protocol (Lu et al., 2007) on 4–11 µm quartz. (B) Dose-response curve for LED 11–355, following SAR protocol (Murray and Wintle, 2000) on 90–180 µm quartz. (C) Natural OSL and IRSL shinedown curves for LED 11–355. Note low IRSL signals, indicating little to no feldspar contamination. Large plot is 4–11 µm quartz, inset shows 90–180 µm quartz. (D) Dose-response curve for LED 11–357, following SMAR protocol (Lu et al., 2007) on 4–11 µm quartz. (E) Dose-response curve for LED 11–357, following SAR protocol (Lu et al., 2007) on 4–11 µm quartz. (E) Dose-response curve for LED 11–357, following SMAR protocol (Lu et al., 2007) on 4–11 µm quartz. (E) Dose-response curve for LED 11–357, following SAR protocol (Murray and Wintle, 2000) on 90–180 µm quartz. Note different scale than LED 11–357 fine-grain sample. (F) Natural OSL and IRSL shinedown curves for LED 11–357. Note low IRSL signals, indicating little to no feldspar contamination. Large plot is 4–11 µm quartz. Additional OSL and IRSL shinedown curves and dose-response curves shown in supplementary material.

Table 2
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Chemistry of OSL samples.

Sample No.	Bulk Alpha (ks-1.cm-2)	U (ppm)	Th (ppm)	K (%)	Rb (ppm)	SWC (%) <sup>a</sup>	WC (%) <sup>a</sup>	Dose rate (Gy/ka)	Dose rate <sup>d</sup> (Gy/ka)
Mayikake									
LED11-210	$5.7 \pm 0.1$	1.4	4.8	0.94	38.6	31	$15 \pm 15$	$2.3 \pm 0.6^{b}$	$1.5 \pm 0.2$
LED11-209	$8.1 \pm 0.2$	2.0	6.8	1.45	62.5	29	$14 \pm 14$	$2.3 \pm 0.6^{b}$	-
LED11-355	$4.9 \pm 0.1$	1.3	4.7	0.98	39.5	20	$10 \pm 10$	$2.1 \pm 0.4^{b}$	$1.6 \pm 0.1$
LED11-356	$5.8 \pm 0.1$	-	-	1.17	-	23	$11 \pm 11$	$2.3 \pm 0.6^{\circ}$	$1.9 \pm 1.3$
Mingyaole									
LED11-360	$6.0 \pm 0.1$	2.0	6.6	1.22	53.3	19	9 ± 9	$2.3 \pm 0.4^{b}$	$2.2 \pm 0.2$
LED11-357	$4.8 \pm 0.1$	1.4	6.0	1.24	53.7	22	$11 \pm 11$	$2.1 \pm 0.4^{b}$	$2.0 \pm 0.2$
LED11-359	$5.3 \pm 0.1$	1.3	4.9	0.95	43.6	23	$11 \pm 11$	$1.7 \pm 0.4^{b}$	-
LED11-358	$5.3 \pm 0.1$	1.5	3.7	0.77	32.6	24	$12 \pm 12$	$1.5 \pm 0.2^{b}$	-

<sup>a</sup> SWC stands for lab-measured saturated water content of the sample. WC stands for water content, defined as weight of water in sample/weight of dry sample. The water content assumes an average of 0% (dry sample) and measured saturated water content.

<sup>b</sup> Fine-grain dose-rate, calculated using U, Th, K, and Rb.

 $^{\rm c}$  Fine-grain dose-rate, calculated using bulk alpha counts (from U and Th) and K.

<sup>d</sup> Coarse-grain dose-rate, calculated using U, Th, K, and Rb, except for LED 11-356, which was calculated using the bulk alpha counts (from U and Th) and K.

**Repka et al., 1997).** The sand samples were collected at intervals of 30–40 cm at depths of 0, 30, 60, 90, 120, 160, and 200 cm below the top of the terrace surface by extracting the sand-sized fraction from the matrix of the gravel cover. Due to the lack of sand at the surface, we also collected a pebble sample from the surface, in addition to pebbles from 2-m depth to constrain the inheritance for the pebbles in comparison to that of sand of the same depositional age. Pebble clasts with diameters of 1–4 cm and high quartz content (primarily granite, vein quartz, and quartzite) were collected (~30 pebbles per depth) and

crushed. Both sand and pebble samples were sieved to  $\,\sim\!0.25\text{--}1$  mm.

We processed the samples at the Cosmogenic Radionuclide Target Preparation Lab at University of California, Santa Barbara, following standard laboratory procedures outlined in the UCSB Cosmogenic Radionuclide Target Preparation Facility Sample Preparation Manual (c.f. Bookhagen and Strecker, 2012). We verified the purity of the quartz using ICP-MS measurements of Al, which yielded concentrations of < 220 ppm in all samples. <sup>10</sup>Be measurements were made at Purdue Rare Isotope Measurement Laboratory (PRIME) Laboratory using the 07KNSTD standard (Nishiizumi et al., 2007).  $^{10}$ Be/ $^9$ Be ratios were corrected using a  $^{10}$ Be laboratory blank (n = 2) of 5.4 × 10<sup>-15</sup> atoms/g for depth-profile sand samples, and 9.2 × 10<sup>-15</sup> atoms/g for pebble samples.

Using the Matlab Monte Carlo modeling program (v. 1.2) from Hidy et al. (2010) and CRONUS Earth 2.2 calculator (Balco et al., 2008), we calculated the age, <sup>10</sup>Be inheritance, and surface erosion rate of each sand-and-pebble depth profile. We calculated cosmogenic <sup>10</sup>Be ages following the constant (time-independent) scaling scheme of Lal (1991) and Stone (2000) and a reference spallogenic <sup>10</sup>Be production rate of 4.01  $\pm$  0.39 atoms/g/a (1 $\sigma$ , Sea Level High Latitude – SLHL) (Borchers et al., 2015) scaled to our field site, a <sup>10</sup>Be half-life of  $1.387 \times 10^6$  years (Korschinek et al., 2010), and an attenuation length of  $160 \text{ g/cm}^2$ (Gosse and Phillips, 2001). We measured topographic shielding in the field (Nishiizumi et al., 1989) and calculated shielding values using the CRONUS Earth 2.2 calculator (Balco et al., 2008). To account for a range of probable overall sediment densities, we applied a density of  $1.5-2.0 \text{ g/cm}^3$  in the Monte Carlo model based on field measurements. All <sup>10</sup>Be depth profile ages and modeled parameters (i.e., inheritance, erosion rate, slip rates) are presented at the 95%-confidence level based on outputs from the Monte Carlo model.

In addition, field observations suggest that little erosion of the terrace surface has occurred since deposition. We did not observe any evidence of significant modification of the terrace surface, such as a fine-grain layer in the subsurface that might indicate inflation, a coarsening of material in the immediate subsurface that may indicate deflation, or the presence of cut-and-fill or other recent fluvial features that may represent erosion of the surface by small channels or overland flow. Therefore, we assigned the maximum surface-erosion depth to be  $\leq 10$  cm and limited the maximum erosion rate of the terrace surface to  $\leq 2$  cm/ka in the Monte Carlo model. This approach of placing limits on erosional parameters based on field evidence is a common practice (Hidy et al., 2010; Haghipour et al., 2012). Below we discuss the uncertainties associated with varying these parameters (section 7.3).

#### 6. Results

In total, 5 coarse-grain and 8 fine-grain quartz OSL samples were measured (Tables 1–3, Figs. 2, Fig. 3, Fig. 4). Coarse-grain and fine-grain samples were taken from the same sampling tube and have the

## Table 3

OSL dating results.

same respective sample numbers. All samples had enough fine-grain material for analysis; however, only 5 samples had enough coarse-grain quartz for analysis. Sixty aliquots were measured for each coarse-grain quartz sample, with the exception of LED 11-356, from which we measured 94 aliquots. Accepted aliquots ranged from 18 for LED 11-210 to 46 for LED 11-360 (Table 3). The  $D_{\rho}$  distributions of the coarse-grain samples (Fig. 4) are skewed, with tails of high  $D_e$  values. Furthermore, well-bleached samples typically have an overdispersion of ~20% (Arnold and Roberts, 2009), whereas all of the coarse-grain quartz samples that we analyzed had overdispersion values ranging from 28 to 58% (Table 3), which may also arise from the use of multiple-grain aliquots and mask the true signal distribution. Based on our  $D_e$  distributions and the observed overdispersion, heterogeneous bleaching appears to be an issue for all of the samples. Thus, the minimum age model (MAM) became our age model of choice (Galbraith et al., 1999), using a  $\sigma_b$  value of 0.1. In section 7.2, we discuss the effect different  $\sigma_b$  values have on the MAM age calculations. We also applied the central age model (CAM) (Table 3) to compare to the fine-grain samples, which were calculated using only the CAM given the low number (n = 10-12) of aliquots measured in the SMAR protocol (Lu et al., 2007). This protocol does not allow for a full evaluation of the  $D_e$ distribution, because the multiple aliquot approach averages interaliquot variations that may have persisted despite the use of multiplegrain aliquots. To calculate ages for each terrace, we take the errorweighted average of all coarse-grain quartz OSL samples following the MAM.

The cosmogenic nuclide depth profiles yielded ages for both sandand pebble-depth profiles on both the Mayikake and Mingyaole terraces.

#### 6.1. Mayikake

At Mayikake, we dated 4 fine-grain and 3 coarse-grain OSL samples, in addition to the cosmogenic<sup>10</sup>Be depth profile that consisted of 6 sand samples and 2 pebble samples. The Mayikake fine-grain samples (n = 4) yield a mean CAM age of 23.2  $\pm$  1.4 ka (Table 3). The 3 coarse-grain quartz samples yielded a CAM age of 25.5  $\pm$  1.8 ka and a MAM age of 16.2  $\pm$  0.8 ka (Table 3).

The cosmogenic <sup>10</sup>Be depth profile based on detrital sand yielded an age of  $14.2^{+3.0}_{-4.4}$  ka (Table 4, Table 5, Fig. 5A). The best-fit, model-derived surface erosion rate for the terrace is 6 mm/ka. The pebble

Sample No.	Aliquots <sup>a</sup>	D <sub>0</sub> (Gy)	Over-dispersion <sup>b</sup>	$CAM^{c} D_{e}$ (Gy)	CAM age (ka)	$\operatorname{MAM}^{\operatorname{d}} D_e$ (Gy)	MAM age (ka)
Mayikake							
LED11-210FQ <sup>e</sup>	10	80	-	$61.5 \pm 2.6$	$21.3 \pm 2.2$	-	-
LED11-209FQ	10	97	-	$61.9 \pm 3.4$	$21.4 \pm 2.5$	-	-
LED11-355FQ	11	67	-	$76.5 \pm 13.8$	$36.4 \pm 7.2$	-	-
LED11-356FQ	11	182	-	$68.3 \pm 7.2$	$29.7 \pm 3.9$	-	-
LED11-210SA <sup>f</sup>	18 (60)	112	$28.0 \pm 4.9$	$42.3 \pm 2.1$	$27.4 \pm 3.5$	$32.1 \pm 2.8$	$20.9 \pm 2.0$
LED11-355SA	19 (60)	152	$40.2 \pm 6.5$	$38.1 \pm 3.5$	$23.7 \pm 3.1$	$21.0 \pm 2.5$	$13.0 \pm 1.6$
LED11-356SA	35 (94)	183	$31.8 \pm 4.0$	$45.4 \pm 2.5$	$25.8 \pm 2.9$	$29.0 \pm 0.5$	$16.5 \pm 1.7$
Mingyaole							
LED11-360FQ	12	217	-	$259.9 \pm 36.4$	$113.0 \pm 18.0$	-	-
LED11-357FQ	12	310	-	$229.5 \pm 22.5$	$111.0 \pm 15.4$	-	-
LED11-359FQ	11	326	-	$215.0 \pm 20.8$	$126.5 \pm 16.9$	-	-
LED11-358FQ	12	240	-	$309.2 \pm 17.9$	$208.5 \pm 31.8$	-	-
LED11-357SA	36 (60)	-	$58.1 \pm 7.0$	$56.3 \pm 6.4$	$28.5 \pm 4.3$	$26.0 \pm 2.7$	$13.2 \pm 1.4$
LED11-360SA	45 (60)	86	$52.0 \pm 5.8$	$44.7 \pm 3.6$	$20.5 \pm 2.4$	$18.6 \pm 1.7$	$8.5 \pm 0.9$

<sup>a</sup> Number of accepted aliquots used in equivalent dose  $(D_e)$  calculations, out of total aliquots measured.

<sup>b</sup> All errors in table are 1 standard error.

<sup>c</sup> CAM – central age model.

<sup>d</sup> MAM – minimum age model. Results show here are MAM-3 model. MAM-4 results are similar but not shown, as p-value was near 0, indicating the MAM-4 model is not a good match for the distribution.

<sup>e</sup> FO – fine-grain quartz.

<sup>f</sup> SA – small-aliquot coarse-grain quartz.



Fig. 4. Coarse-grain quartz OSL sample data. (A–E) Left-hand panels are cumulative frequency (grey circles, with 1 standard error). N is the number of accepted aliquots (out of total aliquots measured). Right-hand panels are radial plots of the same data.  $D_e$  are listed with 1 standard error.

Cosmogenic beryllium-10 data.

Sample No.	Latitude/Longitude	Elevation (m)	Depth (m)	Thickness (cm) <sup>a</sup>	Mass qtz (g)	<sup>10</sup> Be/ <sup>9</sup> Be (10 <sup>-14</sup> ) <sup>bc</sup>	<sup>9</sup> Be carrier (mg) <sup>b</sup>	[ <sup>10</sup> Be] (10 <sup>4</sup> atoms g <sup>-1</sup> )		
Mayikake Depth Profile (detrital sand)										
MYK-6	39.5836°N/75.1046°E	1869	0.3	5	30.18	$31.6 \pm 0.88$	0.252	$17.7 \pm 1.51$		
MYK-5	39.5836°N/75.1046°E	1869	0.6	5	40.98	$30.6 \pm 0.83$	0.252	$12.6 \pm 1.07$		
MYK-4	39.5836°N/75.1046°E	1869	0.9	5	43.64	$23.5 \pm 0.59$	0.249	8.96 ± 0.76		
MYK-3	39.5836°N/75.1046°E	1869	1.2	5	30.69	$12.3 \pm 0.36$	0.250	$6.68 \pm 0.57$		
MYK-2	39.5836°N/75.1046°E	1869	1.6	5	43.88	$12.5 \pm 0.44$	0.245	$4.66 \pm 0.41$		
MYK-1	39.5836°N/75.1046°E	1869	2.0	5	60.76	$12.2 \pm 0.43$	0.251	$3.37 \pm 0.29$		
Pebble Sample	Pebble Samples (1–3 cm)									
MYK-7p	39.5836°N/75.1046°E	1869	0	1.5	95.38	$131.0 \pm 8.39$	0.261	$24.0 \pm 2.48$		
MYK-1p	39.5836°N/75.1046°E	1869	2.0	5	56.09	$12.3 \pm 0.79$	0.267	$3.32 \pm 0.34$		
Mingyaole D	epth Profile (detrital sar	nd)								
MYL-6	39.4838°N/75.3821°E	1670	0.3	5	53.53	$46.9 \pm 0.84$	0.225	$13.2 \pm 1.09$		
MYL-5	39.4838°N/75.3821°E	1670	0.6	5	47.24	$32.7 \pm 0.62$	0.256	$11.8 \pm 0.98$		
MYL-4	39.4838°N/75.3821°E	1670	0.9	5	34.73	$18.2 \pm 0.60$	0.257	$9.0 \pm 0.78$		
MYL-3	39.4838°N/75.3821°E	1670	1.2	5	48.95	$23.9 \pm 0.67$	0.263	$8.6 \pm 0.74$		
MYL-2	39.4838°N/75.3821°E	1670	1.6	5	54.34	$22.0 \pm 0.70$	0.253	$6.85 \pm 0.59$		
MYL-1	39.4838°N/75.3821°E	1670	2.0	5	46.92	$16.8 \pm 0.65$	0.263	$6.3 \pm 0.56$		
Pebble Samples (1–3 cm)										
MYL-7p	39.4838°N/75.3821°E	1670	0	1.5	50.08	$89.0 \pm 3.65$	0.255	$17.4 \pm 1.58$		
MYL-1p	39.4838°N/75.3821°E	1670	2.0	5	57.72	$21.3 \pm 1.17$	0.236	$4.98 \pm 0.49$		

<sup>a</sup> Thickness is the height of the unit (in cm) from which the sample was collected, for sand and pebble samples from depth profiles. Thickness for surface pebble samples refers to the average diameter of a pebble at the surface.

<sup>b</sup> Isotope ratios were normalized to <sup>10</sup>Be standards prepared by Nishiizumi et al. (2007) with a value of  $2.85 \times 10^{12}$  and  $a^{10}$ Be half-life of  $1.387 \times 10^{6}$  years (Korschinek et al., 2010). <sup>c</sup> <sup>10</sup>Be/<sup>9</sup>Be ratios were corrected using  $a^{10}$ Be laboratory blank (n = 2) of  $5.4 \times 10^{-15}$  atoms/g for depth profile sand samples, and  $9.2 \times 10^{-15}$  atoms/g for pebble samples.

samples yielded an equivalent age of  $14.2^{+2.8}_{-4.0}$  ka (Fig. 5A). The inheritances for the sand and pebble fractions were  $0.68^{+1.47}_{-0.68} \times 10^4$  atoms/g and  $0.68^{+1.17}_{-0.68} \times 10^4$  atoms/g, respectively (Table 5).

# 6.2. Mingyaole

Within error, the CRN ages and coarse-grain MAM ages agree. These ages suggest the terrace cover was deposited until  $\sim$  12–14 ka and then abandoned abruptly after deposition (Fig. 6A).

Using the CRN abandonment age for the Mayikake surface, a scarp height and fault dip of 38 m and 75° ± 5° for the Bieertuokuoyi Frontal Thrust (PFT) (Fig. 7A and B) and a scarp height and fault dip of 15 m and 16° ± 3° for the Mayikake Thrust (Fig. 7C and D) that deform this surface, we calculate dip-slip rates of  $2.6^{+0.5}_{-0.5}$  mm/a and  $3.7^{+2.8}_{-1.4}$  mm/a, respectively, for these faults. Given observations on regional seismic lines, we assume the faults merge into a subhorizontal decollement at depth (Chen et al., 2010; Li et al., 2012; Wang et al., 2016). Given the strike-slip component on the Bieertuokuoyi Frontal Thrust described by Li et al. (2012), its estimated total shortening rate is ~4.6 mm/a, whereas for the Mayikake Thrust, the estimated shortening rate is ~3.7 mm/a.

#### Table 5

Cosmogenic beryllium-10 data and ages.

From the Mingyaole terrace, we collected 4 OSL samples, of which we dated all four using the fine-grain component, but only dated two with the coarse-grain component. In addition, we collected a cosmogenic <sup>10</sup>Be depth profile that consisted of 6 sand samples and 2 pebble samples. Three fine-grain samples from the Mingyaole surface yield an average CAM age of 117.8  $\pm$  8.0 ka, with a fourth sample yielding an age of 208.5 ± 24.2 ka (Table 3). Notably, all ages were in stratigraphic order. We interpret that the oldest sample may be near saturation or incompletely bleached, because that sample is nearly 100 ka older than the remaining three samples, yet it lies only  $0.2\text{--}0.4\,\text{m}$ stratigraphically deeper. Moreover, no intervening erosion surfaces or soils were observed in the field or in the cosmogenic depth profile (Fig. 6B). Notably, fine-grain quartz may begin to saturate around 200-300 Gy (Timar-Gabor and Wintle, 2013). Sample LED 11-358 has a  $D_e$  of 309.2 Gy, yet a  $D_0$  of 240 Gy (Table 3), suggesting it is likely saturated. Alternatively, this older age may also be a result of a different dose-rate history (Table 3), or sediment that experienced a

Sample No.	Production Rate (atoms/g/a)		Shielding Correction <sup>c</sup>	Erosion rate (cm/a) <sup>d</sup>	Inheritance (10 <sup>4</sup> atoms/g)	Age (ka) (95% CI) <sup>e</sup>					
	Spallation <sup>a</sup>	Muons <sup>b</sup>									
Ages based on sand samples from 0.3–2 m depth											
Mayikake	16.26	0.330	0.98	0.0006	$0.68^{+1.47}_{-0.68}$	$14.2^{+3.0}_{-4.4}$					
Mingyaole	14.10	0.311	0.98	0.0014	5.09 <sup>+1.82</sup> /-2.37	$8.5^{+4.3}_{-5.0}$					
Ages for pebble	s from 2 m depth an	d the surface									
Mayikake	16.26	0.330	0.99	0.0006	$0.68^{+1.17}_{-0.68}$	$14.2^{+2.8}_{-4.0}$					
Mingyaole	14.10	0.311	0.98	0.0014	$3.21^{+1.28}_{-1.51}$	$10.4^{+2.6}_{-3.5}$					
Ages based on sand and pebble samples											
Mayikake	16.26	0.330	0.98	0.0005	$0.57^{+1.58}_{-0.57}$	$14.2^{+3.0}_{-4.4}$					
Mingyaole	14.10	0.311	0.98	0.0012	4.34 <sup>+1.72</sup> /-2.26	9.4 <sup>+4.2</sup> /-4.5					

<sup>a</sup> Constant (time-invariant) local production rate based on Lal (1991) and Stone (2000). A sea level, high latitude production rate of 4.01 <sup>10</sup>Be atoms/g/a quartz was used (Borchers et al., 2015).

<sup>b</sup> Constant (time-invariant) local production rate based on Heisinger et al. (2002a, b).

<sup>c</sup> Geometric shielding correction for topography calculated with the Cosmic-Ray Produced Nuclide Systematics (CRONUS) Earth online calculator (Balco et al., 2008) version 2.2 (http://hess.ess.washington.edu/).

<sup>d</sup> Erosion rate and inheritance from best fit Monte Carlo model of Hidy et al. (2010).

<sup>e</sup> Bayesian most probable age from Monte Carlo model of Hidy et al. (2010), with 95% upper and lower bounds.



**Fig. 5.** Cosmogenic <sup>10</sup>Be depth profile for A) Mayikake terrace surface and B) Mingyaole terrace surface. Circles represent <sup>10</sup>Be concentration of sand samples, with 1 $\sigma$  uncertainties. Squares represent <sup>10</sup>Be concentrations of pebble samples, with 1 $\sigma$  uncertainties. Solid and dashed black lines are the lines of best fit through the sand and pebble sample data, respectively, with grey lines representing the 95% confidence interval of the line of best fit for the sand samples for Mayikake and Mingyaole calculated using the Monte Carlo model of Hidy et al. (2010). Insets show probability density functions of the age from the Monte Carlo model. Solid black line with dark grey fill represents sand age distribution, dashed black line with light grey fill represents shown as dashed vertical lines.

different transport process, such that fewer grains were bleached during transport in the fluvial system, e.g., eroded from a nearby source at night, as discussed in further detail in the discussion. Regardless, we chose to use the three youngest ages to characterize the fine-grain age (=  $\sim 118$  ka) of the Mingyaole surface (Fig. 6B). On the same Mingyaole terrace, the average coarse-grain quartz CAM and MAM OSL ages are 22.4  $\pm$  2.1 and 9.9  $\pm$  0.8 ka (Fig. 6B, Table 3), respectively.

The cosmogenic <sup>10</sup>Be depth profile based on detrital sand yielded an age of  $8.5^{+4.3}_{-5.0}$  ka (Table 4, Table 5, Fig. 5B). The best-fit, modeled-derived surface erosion rate is 14 mm/ka. The pebbles samples at the surface yielded a similar age of  $10.4^{+2.6}_{-2.5}$  ka (Fig. 5B). The inheritances for the sand and pebble fractions were  $5.09^{+1.82}_{-2.37} \times 10^4$  atoms/g and  $3.21^{+1.28}_{-4.051} \times 10^4$  atoms/g, respectively (Table 5).

The CRN ages and coarse-grain MAM ages agree well. These ages suggest the terrace cover was deposited until  $\sim$ 8–10 ka and then abandoned shortly thereafter: approximately 8.5 ka (Fig. 6B).

Using the CRN sand depth-profile age of ~8.5 ka and an incremental shortening of ~10.1 m, we calculate a shortening rate of  $1.2^{+3.9}_{-0.3}$ mm/a for the southern limb of the Mingyaole anticline following the equations outlined in Li et al. (2015b) (Fig. 7E and F). If we assume the northern limb is shortening at the same rate (Scharer et al., 2006), we calculate an overall shortening rate of ~2.4 mm/a for the Mingyaole anticline since the terrace abandonment age of ~8.5 ka.

#### 7. Discussion

#### 7.1. OSL grain size-dependent age differences

In the western Tarim Basin, we observe that fine-grain OSL samples likely overestimate the depositional age of a terrace (Fig. 6). On Mayikake, the fine-grain CAM ages agree with the coarse-grain CAM ages within error. In contrast, the fine-grain CAM ages from the



Fig. 6. Age-depth profiles for each site, comparing OSL ages calculated using the two different age models (discussed in text) and cosmogenic <sup>10</sup>Be sand and pebble depth-profile ages. (A) Mayikake terrace. (B) Mingyaole terrace. CAM – central age model, MAM – minimum age model. OSL error bars are 1 standard error, <sup>10</sup>Be error bars are 95% confidence intervals.



Fig. 7. (A) dGPS topographic profile and (B) field photo of the Bieertuokuoyi Frontal Thrust on the Mayikake surface. Modified from Li et al. (2012). (C) dGPS topographic profile and (D) field photo of the Mayikake Thrust on the Mayikake surface. Modified from Li et al. (2012). (E) dGPS topographic profile of the T2 surface and (F) field photo of the T3b surface of the Mingyaole fold scarp. Modified from Li et al. (2015b). See Fig. 2 for locations of dGPS profiles and field photos.

Mingyaole surface overestimate the coarse-grain CAM age by  $\sim 100$  ka (a factor of ten). This discrepancy may arise for several reasons: (1) different aliquot sizes, leading to a different number of grains per disc (Wallinga, 2002; Duller, 2008); (2) insufficient bleaching of the finer grains, leading to the retention of residual doses that result in an overestimation of ages; (3) differences in the sources of the silt and sand grain-size fractions, leading to intrinsic differences in the way the quartz behaves. Below, we discuss the aliquot size and observations on the bleaching of grains in the western Tarim Basin. We do not have the required data to discuss different sources of the fine-vs. coarse-grains, but we suggest this potential contrast as an avenue for future research in this region.

## 7.1.1. Aliquot size

Apart from the applied SMAR protocol, the apparent overestimation of the fine-grain OSL ages may result from the different aliquot sizes used for fine- and coarse-grain dating (diameters of 9.7 mm for finegrain, and 2 mm for coarse-grain). The differences in both aliquot size and grain size result in large differences both in the number of grains on each disc and in the grains that contribute to the luminescence signal. Almost one million grains are present on each fine-grain aliquot: an abundance that averages out any grain-to-grain dose variations (Wallinga, 2002; Duller, 2008), but that can also result in an overestimation of the true depositional age of the sediment. Furthermore, the aliquot size effect is worsened by the use of the SMAR protocol, which not only averages more grains per aliquot, but also averages interaliquot variation. The small-aliquot, coarse-grain quartz samples (90–180  $\mu$ m) have ~ 200–300 grains per disc, and only 1.5–3.6% of the grains from nearby sites emit a luminescence signal (Yang et al., 2017). In homogeneously bleached depositional settings, averaging the dose variations has little effect on the age, but an apparent dose overestimation typifies heterogeneously bleached depositional settings (Arnold and Roberts, 2009; Cunningham et al., 2011).

# 7.1.2. Bleaching of different grain sizes

In semi-arid to arid settings, most of the terrestrial sediment transport and deposition occurs during storms, which can have short, high-flow durations of a few hours, e.g., Porat et al. (2001) and Cohen and Laronne (2005). Sediment-laden flash-flood waters are turbid and may not occur during daylight hours, which may limit penetration of any sunlight and inhibit bleaching of the grains (Berger and Luternauer, 1987; Sanderson et al., 2007; Gray and Mahan, 2014). As a result, fine grains traveling as suspended load may not be exposed to sunlight long enough to be fully bleached prior to deposition (Sanderson et al., 2007). In such cases, most bleaching likely occurs during transient storage in bars along the river, instead of during transport (Gray and Mahan, 2014). Furthermore, fine-grains tend to flocculate and form aggregates, possibly due to mud coatings, which further hinders solar bleaching even when episodic flows occur during daylight (Rittenour, 2008; Hu et al., 2010; Gray and Mahan, 2014). In the western Tarim Basin, fine-

grain sediment in the rivers is commonly transported during episodic, short-lived floods that may erode proximal older geomorphic surfaces and weakly consolidated, exhumed Tertiary bedrock.

Coarser sediment, such as fine sand, is more likely to have been transported as saltating bedload and to spend a longer time exposed on channel bars between floods, especially if traveling as a continuous flow in waters with lower sediment concentration (Gray and Mahan, 2014). We observe that active bars in the Tarim channels are primarily sands. Recent work by Cunningham et al. (2015) tentatively suggests that coarse grains on bars or near the channel edge are likely to be better bleached because of the opportunities for reworking near the surface. Thus, coarse grains deposited repeatedly on exposed bars are likely to get reworked by both wind and water and to be thoroughly bleached as they travel through the fluvial system.

We speculate that river characteristics, especially in arid settings, likely have a dominant effect on the bleaching of fluvial sediments. In our study, OSL data from the Mayikake terrace - a terrace likely deposited by the Bieertuokuoyi River - indicate the sediments are more thoroughly bleached, with overdispersion varying between 28 and 40% and fine-grain and coarse-grain CAM ages in general agreement. The Bieertuokuoyi River is generally shallow and has clear water when not in flood. In contrast, the Kezilesu River (meaning "Red River"), which deposited the Mingyaole terrace, has a typical red-brown color due to its high suspended-sediment load and deeper water depths. The OSL data from the Mingyaole terrace display both higher overdispersion values of 50-58% and a larger discrepancy between the fine- and coarse-grain CAM ages. We conclude that the grains are heterogeneously bleached. Although the current river characteristics may not be indicative of the river characteristics over the last 100 ka, the OSL data indicate that, in general, deposits from the Bieertuokuoyi River are better bleached (sediments in the Mayikake Basin, this study; Li et al., 2012) than deposits from the Kezilesu River, e.g., sediments near the Mingyaole anticline (this study; Li et al., 2015b) and Mushi anticline (Li et al., 2013).

These observations and data, although based on a small sample set, suggest that rivers with higher suspended sediment load and deeper water depths are less likely to fully bleach sediments during transport (Berger, 1990; Berger and Leternauer, 1987; Gray and Mahan, 2014), especially the fine-grain fraction that is more likely to be traveling as suspended sediment during sediment-laden flash floods in semi-arid to arid settings.

#### 7.2. OSL MAM age uncertainties arising from choice of $\sigma_b$

The  $\sigma_b$  parameter of the MAM is defined as the overdispersion of a well-bleached natural sample from a given field area. In the absence of direct data, Galbraith and Roberts (2012) recommend using overdispersion values of 10–20%, values that are approximately equal to many of the  $\sigma_b$  values for individual aliquots and quartz grains in the literature (Arnold and Roberts, 2009). Given that we do not have a well-bleached sample from our field area from which to calculate  $\sigma_b$ , we chose 0.1 because this value represents the overdispersion of the first regenerated dose for all the accepted aliquots. This value should be representative of the intrinsic overdispersion of the quartz that we are able to detect from our samples. Moreover, we also note that our choice of  $\sigma_b$  is similar to other studies on alluvial and fluvial sediments in arid/semi-arid settings, which use a value of ~0.1 (0.1–0.15) (Fattahi et al., 2010; Colarossi et al., 2015).

However, given that we cannot directly calculate  $\sigma_b$ , we tested the sensitivity of our ages to our choice of  $\sigma_b$  by re-running the MAM calculations using  $\sigma_b$  values of 0.1, 0.2, 0.3, and 0.4 (Table S3, Table S4), which are common values for fluvial sediments in the literature (Fattahi et al., 2010; Trauerstein et al., 2014; Colarossi et al., 2015; Wang et al., 2015). Using a  $\sigma_b$  of 0.4 yields MAM  $D_e$ 's that are ~10–20 Gy higher (Table S3) and MAM ages that are ~5–10 ka older (Table S4) than using a  $\sigma_b$  of 0.1, with the exception of LED 11-356, which has an age

that is ~20 ka older. Using a  $\sigma_b$  of 0.2 results in MAM  $D_e$ 's that are, on average, ~2–4 Gy higher (Table S3) and MAM ages that are ~1–3 ka older than using a  $\sigma_b$  of 0.1. A larger  $\sigma_b$  value of up to 0.4 increases the mean age of the Mayikake and Mingyaole surfaces up to 24.9 ± 2.2 and 15.2 ± 2.1 ka, respectively. Importantly, changing the MAM ages and mean ages for the surfaces by ~1–10 ka does not significantly change the tectonic and climatic implications of this study: the derived slip rates rely on the better-constrained CRN abandonment ages on Mingyaole and Mayikake, and the ages still broadly correlate with climatic events, within error (see section 8.2). However, if a larger  $\sigma_b$ value of 0.4 were used, the coarse-grain OSL MAM and CAM ages from Mayikake would agree, but the fine-grain OSL CAM ages on Mayikake and all OSL CAM ages on Mingyaole are still older (with the exception of LED11-210). Based on our above reasoning, we suggest that a  $\sigma_b$  of 0.1 or 0.2 is acceptable for our dataset.

## 7.3. Cosmogenic<sup>10</sup>Be age uncertainties

The age we calculate from the Monte Carlo model (Hidy et al., 2010) is sensitive both to the surface erosion rate since deposition and to the limits placed on the total, post-depositional erosion magnitude. Importantly, if the erosion rate and erosion magnitude are allowed to vary in an unconstrained manner, the Monte Carlo model yields a bestfit erosion rate of 2–3 cm/ka, a total erosion of  $\sim$  20–40 cm, and an age that is 1-3 ka older than if the model uses a more constrained erosion rate and magnitude (as described in Section 5). Notably, these ages with unconstrained errors are still within the 95%-confidence intervals of the ages calculated with constrained parameters. Because field observations indicate no significant erosion, deflation, or inflation of the surface, we use ages calculated with a maximum erosion rate of 2 cm/ka, and a maximum total erosion of 10 cm. The best-fit erosion rate from the Monte Carlo Model was 0.6 cm/ka for Mavikake and 1.4 cm/ka for Mingyaole. However, we acknowledge that a wetter climate during the late Pleistocene and early-mid Holocene (Yang and Scuderi, 2010) might have created conditions with higher erosion rates for a period of time soon after the terraces were deposited and abandoned. Importantly, exponential trends of the concentrations with depth are consistent with continuous and rapid initial aggradation.

## 8. Regional implications

#### 8.1. Deformation rates

The deformation rates calculated on the three structures reveal rapid, shortening during the late Quaternary across the NE Pamir margin (Li et al., 2012) (Fig. 7). Furthermore, the shortening rate across the NE Pamir margin since 0.35 Ma appears to be steady (Li et al., 2012), even though the deformation appears to be accommodated on different structures through time. Over timescales of  $\sim 98$  ka, the shortening rate averages  $> \sim 5.6$  mm/a on the Pamir Frontal Thrust (Thompson Jobe et al., 2017). On a shorter timescale (since the Last Glacial Maximum), the shortening rate ( $\sim 8 \text{ mm/a}$ ) is similar across the same region of the northern margin of the Pamir, but is accommodated on two structures: the Bieertuokuoyi Frontal Thrust (~4.6 mm/a) and the Mayikake Thrust ( $\sim$  3.7 mm/a). These rates are similar to the rate since 0.35 Ma of  $\sim$  6.4 mm/a (Li et al., 2012), and the modern geodetic shortening rate of 6-9 mm/a (Zubovich et al., 2010; Li et al., 2012). Thus, these new results support the idea that the NE Pamir margin has experienced a relatively uniform shortening rate of 6-9 mm/a since 0.35 Ma (Li et al., 2012; Thompson Jobe et al., 2017). The deformation may be accommodated on different structures through time, as shortening shifts spatially from the Pamir Frontal Thrust ~98 ka to being partitioned on both the Bieertuokuoyi Frontal Thrust and the Mayikake Thrust since ~15 ka (Thompson Jobe et al., 2017).



**Fig. 8.** Regional terrace correlations. Published terrace OSL ages and ages from this study shown with Vostok ice-core record from Petit et al. (1999). Glacial periods highlighted with light grey box. OSL ages shown as blue circles with 1 standard error uncertainty. CRN ages are shown as blue squares with 95% confidence interval. Approximate timing of terrace deposition and abandonment, based on OSL and CRN ages, shown as blue bars, and commonly align with glacial-interglacial transitions. OSL ages for Mayikake – Li et al., 2012, this study; Mushi – Li et al., 2013; Mingyaole – Li et al., 2015a, Li et al., 2015b, this study. (For interpretation of the references to color in this figure legend, the reader is referred to the Web version of this article.)

#### 8.2. Regional terrace correlations

Despite spatially varying rates of deformation, the ages on the Mayikake terrace are in general agreement with dated terraces in the western Tarim Basin and along the margins of the Tian Shan: a consistency suggesting a climatic control on the formation of the regional terraces (Fig. 8) (Pan et al., 2003). Nearby terraces at Mushi anticline (Li et al., 2013) and Mingyaole anticline (higher terrace level than in this present study; Li et al., 2015b) have been dated to  $\sim$  15–19 ka using fine-grain and coarse-grain quartz OSL. Terraces on the flanks of the southern (Hubert-Ferrari et al., 2005) and northern (Thompson et al., 2002; Charreau et al., 2011) Tian Shan date to ~13-18 ka. The deposition of the Mayikake surface dates to  $\sim 12-14$  ka. These ages (~12-19 ka) suggest the deposition of regionally extensive terrace surfaces during the last deglaciation (MIS 2 to MIS 1) (Fig. 8), followed by rapid subsequent abandonment. Notably, all of these terraces are crossing active structures. Despite the consistent age of regional terrace deposition, the heights of the terraces above the modern rivers vary, suggesting a primarily climatic control on terrace formation, instead of incision due to active tectonics (Pan et al., 2003; Li et al., 2013; Huang et al., 2014).

The age for the Mingyaole terrace ( $\sim$ 8–10 ka) does not correlate to other known regional terraces (Fig. 8). Yang and Scuderi (2010) determined a complicated climate history from lakebeds, river terraces, and loess for the Holocene in the western Tarim Basin. Approximately 8 ka, a transition occurred from a drier climate in the early Holocene to a wetter climate during the mid-Holocene, but the records show short-lived periods of arid climate during this interval as well (Yang and Scuderi, 2010). This transition at ~8 ka approximately matches the formation and abandonment of the Mingyaole T2 terrace. Alternatively, the end of deposition on this terrace might be a response to accelerated rock uplift due to outward hinge migration on the southern limb of the growing Mingyaole anticline (Li et al., 2015b).

#### 9. Conclusions

We dated deformed fluvial terraces crossing active structures on the NE Pamir margin, northwest China, using OSL and cosmogenic <sup>10</sup>Be dating techniques. Fine-grain (4-11 µm) quartz OSL samples, when measured using the SMAR protocol, may overestimate the age of a terrace up to a factor of ten. Importantly, we find residual doses still present in both fine and coarse (90-180 µm) grain sizes on terraces of Late Pleistocene age: ages that can lead to significant overestimation (as much as an order of magnitude) of the terrace age. Thus, dating approaches need to include an adequate assessment of the dose distributions, i.e., small aliquots consisting of few grains and single-grain measurement protocols. Based on the equivalent dose distribution and stratigraphic consistency of ages, we applied the minimum age model (MAM) with a  $\sigma_b$  of 0.1 to the coarse-grain samples to characterize the depositional ages of the terraces. Applying different  $\sigma_b$  values up to 0.4 may increase the OSL MAM ages by 1-10 ka. The cosmogenic depth profiles yielded predictable depth-dependent concentrations for the Mayikake and Mingyaole terraces, with good agreement between the pebble and sand fractions. Cosmogenic <sup>10</sup>Be depth profile ages from three sites date the abandonment of the terrace surfaces to  $\sim 8$  and  $\sim 14$ ka. The coarse-grain OSL MAM ages date the deposition of the Mingyaole and Mayikake surfaces to  $\sim 9$  and  $\sim 16$  ka, respectively.

Cosmogenic <sup>10</sup>Be and coarse-grain OSL MAM ages agree well. The age of the Mayikake terrace is consistent with other terraces in the western Tarim Basin that date their deposition and subsequent abandonment to the last deglaciation (12–18 ka) (Thompson et al., 2002; Hubert-Ferrari et al., 2005; Li et al., 2012, 2013, 2015b) and suggest the formation of these terraces on the margins of the Tarim Basin and along the flanks of the Tian Shan is climatically controlled (Pan et al., 2003).

Furthermore, new estimated shortening rates of  $\sim 3.7$ ,  $\sim 4.6$ , and  $\sim 2.4$  mm/a on the Mayikake Thrust, Pamir Frontal Thrust, and Mingyaole anticline, respectively, when combined with other published data in the region, suggest temporally uniform shortening across the NE Pamir margin during the Late Quaternary and illuminate the spatial migration of deformation to structures along the Pamir-Tian Shan interface over late Quaternary timescales.

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## Appendix A. Supplementary data

Supplementary data related to this article can be found at http://dx. doi.org/10.1016/j.quageo.2018.01.002.

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